



Moments of Vlasov-Poisson eq.

## A journey from **Antonov** to **Morrison**

B. Després

Laboratoire Jacques-Louis Lions

UPMC-Paris VI, CNRS UMR 7598, Paris, France

September 6, 2013

Introduction

Moments

Stepping from Antonov to Morrison

Influence of the Debye length

Real non linear

Conclusion

A question about magnetic equation

Model problem in dimension 1+1:  $x \in I = [0, 1]_{\text{per}}$  and  $v \in \mathbb{R}$ ,

$$\mathbf{V.P.} \quad \begin{cases} \partial_t f + v \partial_x f - E \partial_v f = 0, & t > 0, \quad (x, v) \in I \times \mathbb{R}, \\ \partial_x E = \sqrt{2\pi} - \int_{\mathbb{R}} f dv, & t > 0, \quad x \in I. \end{cases}$$

Gauss equation is compatible with Ampère equation  $\partial_t E = \int v f dv$  if the total momentum vanishes  $\int_I \int_{\mathbb{R}} f v dv = 0$  at  $t = 0$ . So

$$\mathbf{V.A.} \quad \begin{cases} \partial_t f + v \partial_x f - E \partial_v f = 0, & t > 0, \quad (x, v) \in I \times \mathbb{R}, \\ \partial_t E = \int_{\mathbb{R}} v f dv, & t > 0, \quad x \in I. \end{cases}$$

is the formulation used hereafter. Basic properties are

$$\begin{cases} \frac{d}{dt} \left[ \int_I \int_{\mathbb{R}} f(t, x, v) v^2 dv dx + \int_I E(t, x, v)^2 dx \right] = 0, \\ 0 \leq f(t, x, v) \leq \|f_0\|_{L^\infty(I \times \mathbb{R})}, \quad f(0, x, v) = f_0(x, v). \end{cases}$$

The initial data of **V.A.** always satisfies the Gauss law.

## Introduction

### Moments

### Stepping from Antonov to Morrison

### Influence of the Debye length

### Real non linear

### Conclusion

### A question about magnetic equation

We consider  $f = G + g$ , with  $G(v) = \exp(-\frac{v^2}{2})$ , so that  $\int G(v)dv = \sqrt{2\pi}$  and the electric field is already first order. Linearized equations are

$$\begin{cases} \partial_t g + v \partial_x g &= -EvG, & t > 0, & (x, v) \in I \times \mathbb{R}, \\ \partial_t E &= \int_{\mathbb{R}} vg dv, & t > 0, & x \in I. \end{cases}$$

The **Antonov** energy identity writes

$$\frac{d}{dt} \left( \int_I \int_{\mathbb{R}} \frac{g^2}{G} dv dx + \int_I E^2 dx \right) = 0$$

In gravitational equations there is actually a minus:  $-\int_I E^2$ .

- Antonov: Remarks on the problem of stability in stellar dynamics, 1961
- Lemou-Mehats-Raphael:  $\approx$  "Variational stability of gravitational systems", 2011
- Mouhot: Stabilité orbitale pour le système de Vlasov-Poisson gravitationnel, 2012

## Introduction

## Moments

Stepping from  
Antonov to  
MorrisonInfluence of the  
Debye length

## Real non linear

## Conclusion

A question about  
magnetic  
equation

Set  $M(v) = \sqrt{G(v)} = \exp(-\frac{v^2}{4})$ .

The "good" unknown is:  $(u, E) = \left(\frac{g}{M(v)}, E\right)$

$$\begin{cases} \partial_t u + v \partial_x u &= -vME, & t > 0, & (x, v) \in I \times \mathbb{R}, \\ \partial_t E &= \int_{\mathbb{R}} uvMdv, & t > 0, & x \in I, \end{cases}$$

with the energy identity  $\frac{d}{dt} \left( \int_I \int_{\mathbb{R}} u^2 dv dx + \int_I E^2 dx \right) = 0$ .

First goal is to analyze the structure of the underlying isometry group

$$(u_0, E_0) \longrightarrow \mathcal{L}(t)(u_0, E_0) = (u(t), E(t)).$$

- Kato, Perturbation theory for linear operators, 1966.

# Hermite polynomials

Introduction

Moments

Stepping from  
Antonov to  
Morrison

Influence of the  
Debye length

Real non linear

Conclusion

A question about  
magnetic  
equation

- Hilbert basis of the space  $\int_{\mathbb{R}} f^2(v)G(v)dv < \infty$ .
  - Rodrigue's representation  $H_n(v) = (-1)^n G(v)^{-1} \frac{d^n}{dv^n} G(v)$ .
  - The degree of  $H_n$  is  $n$ . The parity of  $H_n$  is the parity of  $n$ .
  - $\int H_n(v)H_m(v)G(v)dv = (2\pi)^{\frac{1}{2}} n! \delta_{nm}$ ,  $n, m \in \mathbb{N}$ .
  - Recursion formula

$$H_{n+1}(v) = vH_n(v) - H_{n-1}(v).$$

- Hermite functions  $\psi_n(v) = (2\pi)^{-\frac{1}{4}} n!^{-\frac{1}{2}} H_n(v)M(v)$  are more convenient:

$$\psi_0(v) = \frac{M(v)}{(2\pi)^{\frac{1}{4}}}, \quad \psi_1(v) = \frac{vM(v)}{(2\pi)^{\frac{1}{4}}}, \quad \psi_2(v) = \frac{(v^2 - 1)M(v)}{(8\pi)^{\frac{1}{4}}}, \quad \dots$$

$$\text{Recursion: } v\psi_n(v) = \sqrt{n+1}\psi_{n+1}(v) + \sqrt{n}\psi_{n-1}(v).$$

- Hammett-Dorland-Perkins: Fluid models of phase mixing, Landau damping, and nonlinear gyrokinetics dynamics, 1992

# Writing the moments

Introduction

**Moments**

Stepping from  
Antonov to  
Morrison

Influence of the  
Debye length

Real non linear

Conclusion

A question about  
magnetic  
equation

The moments of  $u$  are:  $\alpha_n = \int_{\mathbb{R}} u \psi_n dv$

$$u(t, x, v) = \sum_{n \in \mathbb{N}} \alpha_n(t, x) \psi_n(v) \text{ with } \|u\|_{L^2(I \times \mathbb{R})}^2 = \sum_{n \in \mathbb{N}} \|\alpha_n\|_{L^2(I)}^2.$$

Taking the moments of the evolution equation, the equations are now

$$\begin{cases} \partial_t \alpha_0 + \sqrt{n+1} \partial_x \alpha_{n+1} + \sqrt{n} \partial_x \alpha_{n-1} = -(2\pi)^{\frac{1}{4}} \delta_{n,1} E, & n = 0, 1, 2, \dots \\ \partial_t E = (2\pi)^{\frac{1}{4}} \alpha_1. \end{cases}$$

# Infinite linear system (infinite wave equation)

Introduction

Moments

Stepping from Antonov to Morrison

Influence of the Debye length

Real non linear

Conclusion

A question about magnetic equation

The infinite vector  $U = \begin{pmatrix} E \\ \alpha_0 \\ \alpha_1 \\ \dots \end{pmatrix}$  satisfies:  $\partial_t U + A \partial_x U = -iBU$

where the matrices are very sparse

$$A = \left( \begin{array}{ccc|ccc} 0 & 0 & 0 & 0 & 0 & \dots \\ 0 & 0 & 1 & 0 & 0 & \dots \\ 0 & 1 & 0 & 2^{\frac{1}{2}} & 0 & \dots \\ \hline 0 & 0 & 2^{\frac{1}{2}} & 0 & 3^{\frac{1}{2}} & 0 \\ 0 & 0 & 0 & 3^{\frac{1}{2}} & 0 & 4^{\frac{1}{2}} \\ \dots & \dots & \dots & \dots & 4^{\frac{1}{2}} & \dots \end{array} \right)$$

and  $B$  is finite rank

$$B = \left( \begin{array}{ccc|ccc} 0 & 0 & -i(2\pi)^{\frac{1}{4}} & 0 & 0 & \dots \\ 0 & 0 & 0 & 0 & 0 & \dots \\ i(2\pi)^{\frac{1}{4}} & 0 & 0 & 0 & 0 & \dots \\ \hline 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 \\ \dots & \dots & \dots & \dots & \dots & \dots \end{array} \right) \cdot$$

From  $\hat{U}_k = \hat{U}(k) = \int_I e^{-ikx} U(x) dx$  for  $k \in \mathbb{Z}$ , one gets mode per mode

$$\partial_t \hat{U}(k) = -iA_k \hat{U}(k)$$

where  $A_k$  is an infinite matrix

$$A_k = \left( \begin{array}{ccc|ccc} 0 & 0 & -i(2\pi)^{\frac{1}{4}} & 0 & 0 & \dots \\ 0 & 0 & k & 0 & 0 & \dots \\ i(2\pi)^{\frac{1}{4}} & k & 0 & k2^{\frac{1}{2}} & 0 & \dots \\ \hline 0 & 0 & k2^{\frac{1}{2}} & 0 & k3^{\frac{1}{2}} & 0 \\ 0 & 0 & 0 & k3^{\frac{1}{2}} & 0 & k4^{\frac{1}{2}} \\ \dots & \dots & \dots & \dots & k4^{\frac{1}{2}} & \dots \end{array} \right)$$

Notice that  $A_k$  an unbounded auto-adjoint operator in  $l_2 \equiv \{W = (w_i)_{i \in \mathbb{N}}, \sum_{i \in \mathbb{N}} |w_i|^2 < \infty\}$ .

Scattering theory says: compute the spectrum of  $kA$ , then add the finite rank perturbation  $-iB$ .

- Kato, Perturbation theory for linear operators, 1966.



The spectrum is made of two parts

$$S(A) = S_d(A) \cup S_c(A)$$

where

- $S_d(A) = \{0\}$  is pure spectrum with eigenvector is  $\mathbf{e}_1 = (1, 0, 0, \dots)^t$ .
- $S_c(A) = \mathbf{R}$  is continuous spectrum: generalized eigenvectors are  $U_\lambda$  for  $\lambda \in \mathbf{R}$ :

$$U_\lambda = \begin{pmatrix} 0 \\ \psi_0(\lambda) \\ \psi_1(\lambda) \\ \psi_2(\lambda) \\ \dots \end{pmatrix} \notin l_2.$$

The proof is by the recursion formula.

# Spectrum of $A_k$ , $k \neq 0$

Introduction

Moments

Stepping from  
Antonov to  
Morrison

Influence of the  
Debye length

Real non linear

Conclusion

A question about  
magnetic  
equation

The spectrum of  $A_k$ ,  $k \neq 0$ , is

$$S(A_k) = S_d(A_k) \cup S_c(A_k)$$

- $S_d(A_k) = \{0\}$  : the eigenvector is  $v_0^k = (k, -i\alpha, 0, 0, \dots)^t \in l^2$ .
- $S_c(A_k) = \mathbb{R}$ : the generalized eigenvector associated to the eigenvalue  $\lambda k$  is

$$\begin{aligned} V_\lambda^k &= -iM(\lambda)ke_1 \\ &+ \left( k^2 + P.V. \int_{\mathbb{R}} \frac{\mu}{\mu - \lambda} M(\mu)^2 d\mu \right) U_\lambda \\ &- M(\lambda)P.V. \int_{\mathbb{R}} \frac{\mu}{\mu - \lambda} U_\mu M(\mu) d\mu \end{aligned}$$

where  $P.V.$  indicates principal value.

- Stein: Singular integrals and differentiability of functions, Princeton  
Mathematical Series, 1970

# Where does this strange formula come from?

Introduction

Moments

Stepping from Antonov to Morrison

Influence of the Debye length

Real non linear

Conclusion

A question about magnetic equation

Search for a formula

$$V_\lambda = a_\lambda \mathbf{e}_1 + b_\lambda U_\lambda + P.V. \int_{\mathbb{R}} \frac{c_\lambda(\mu)}{\mu - \lambda} U_\mu d\mu$$

where  $a_\lambda$ ,  $b_\lambda$  and  $c_\lambda(\mu)$  are unknown quantities.

The number of unknowns in this equation is "infinite" due to the unknown function  $c_\lambda$ .

The eigenvector equation  $A_k V_\lambda = k\lambda V_\lambda$  is equivalent to

$$(A_k V_\lambda, \mathbf{e}_n) = k\lambda (V_\lambda, \mathbf{e}_n), \quad \forall n \in \mathbb{N}$$

with

$$A_k = kA - iB$$

a sparse "tridiagonal" hermitian matrix with all coefficients known.

Trick: for all  $n \geq P \gg \text{rank}(B)$  it reduces to

$$\begin{aligned} & (k\lambda U_\lambda, \mathbf{e}_n) + P.V. \int_{\mathbb{R}} \frac{c_\lambda(\mu)}{\mu - \lambda} (k\mu U_\mu, \mathbf{e}_n) d\mu \\ &= (k\lambda U_\lambda, \mathbf{e}_n) + P.V. \int_{\mathbb{R}} \frac{c_\lambda(\mu)}{\mu - \lambda} (k\lambda U_\mu, \mathbf{e}_n) d\mu. \\ &\iff kP.V. \int_{\mathbb{R}} c_\lambda(\mu) \psi_n(\mu) d\mu = 0, \quad \forall n \geq P \\ &\iff c_\lambda \in \text{Span}(\psi_m)_{m < P}. \end{aligned}$$

Therefore one has to deal only with a finite number of coefficients!!  
The solution is after that easy to compute.

The form of the final formula is (of course) very similar to the method of normal modes of Van Kampen

- Van Kampen: On the theory of stationary waves in plasmas, 1955
- K.M. Case, Plasma oscillations, Annals of Physics, 1959.
- Morrison and Pfirsch: Dielectric energy versus plasma energy, and Hamiltonian action-angle variables for the Vlasov equation, 1992.

# Integro-differential operator

Introduction

Moments

Stepping from Antonov to Morrison

Influence of the Debye length

Real non linear

Conclusion

A question about magnetic equation

- The formula for the  $(V_\lambda)$  in terms of the  $(U_\lambda)$  is a kind of change of the "orthogonal" basis.

- Its enough to write directly its effect on  $f(\lambda) = (W, U_\lambda)$  for an arbitrary  $W \in l_2$ . This is done with an integro-differential operator.

- The integro-differential operator writes

$$Lu = (-\partial_{xx} + q(v)) u(x, v) - vM(v)P.V. \int_{\mathbb{R}} \frac{1}{w - v} u(x, w) M(w) dw$$

where the function  $q$  is defined by

$$q(v) = P.V. \int_{\mathbb{R}} \frac{w}{w - v} M(w)^2 dw.$$

Up to the use of  $u = \frac{g}{M}$  instead of  $g = f - M^2$  it is the same  $\mathcal{G}$  as in  
- Morrison: Hamiltonian Description of Vlasov Dynamics: Action-Angle Variables for the Continuous Spectrum, 2000.

# Main magic property

Introduction

Moments

Stepping from  
Antonov to  
Morrison

Influence of the  
Debye length

Real non linear

Conclusion

A question about  
magnetic  
equation

Take  $u$  that satisfies  $\partial_t u + v\partial_x u + vME = 0$  with the Gauss law.  
Then  $h = Lu$  is a solution of free transport

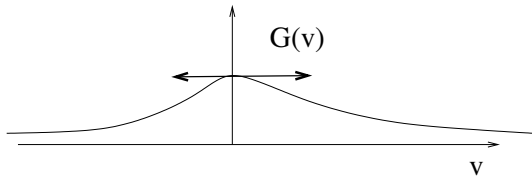
$$\partial_t h + v\partial_x h = 0.$$

Direct proof is as follows

$$\begin{array}{rcl} L\partial_t u & = & \partial_t h \\ L(v\partial_x u) & = & v\partial_x Lu - vM(v)\partial_x \int_{\mathbb{R}} u(x, w)M(w)dw \\ L(vM(v)E) & = & -vM(v)\partial_{xx} E \end{array}$$

The third column collapses thanks to the Gauss law.

- The operator  $u \mapsto q(v)u(x, v) - vM(v)P.V. \int_{\mathbb{R}} \frac{1}{w-v} u(x, w)M(w)dw$  is continuous, Hilbert-Schmidt, triangular in the Hermite basis with almost zero diagonal.
- $L$  is invertible: boundedness of  $L^{-1}$  in  $l_2$  Fourier mode per Fourier mode,
- Exact formula for  $L^{-1}$  (Morrison, ...).
- The moment part of the proof is immediately generalizable to "one bump" equilibrium



since  $M^2(v) = -\frac{G(v)}{v} > 0$  is still positive.

# Linear Landau damping with low regularity

Introduction

Moments

Stepping from Antonov to Morrison

Influence of the Debye length

Real non linear

Conclusion

A question about magnetic equation

- Consider a solution  $(u, E)$  of the linearized **V.P.** equations. Assume the initial condition is  $u_0 \in L^2(I : H^n(\mathbb{R}))$  with  $n \geq 1$ . There exists a constant  $C_n > 0$  such that

$$\|\phi(t)\|_{W^{1,\infty}(I)} + \|\partial_t \phi(t)\|_{L^\infty(I)} \leq \frac{C_n}{\max(1, t^n)} \|u_0\|_{L^2(I; H^n(\mathbb{R}))}.$$

where  $\phi$  is the electric potential

$$-\partial_x \phi = E, \quad \int_I \phi dx = 0.$$

- Key mechanism  $\int_{\mathbb{R}} h_0(x - vt) \varphi(v) dv \approx \sum_{k \neq 0} \int_{\mathbb{R}} e^{-ikvt} f_k(v) dv$  with

$$\int_{\mathbb{R}} e^{-ikvt} f_k(v) dv = \frac{1}{(-ikt)^n} \int_{\mathbb{R}} e^{-ikvt} \frac{d^n}{dv^n} f_k(v) dv.$$

Classical proof *à la Landau* in

- Sonnendrucker: Modèles cinétiques pour la fusion, 2008.
- Mouhot and C. Villani, On Landau damping, 2011.



# V. plus non linear P. equation

Introduction

Moments

Stepping from Antonov to Morrison

Influence of the Debye length

Real non linear

Conclusion

A question about magnetic equation

Non linear Vlasov-Poisson equation for ions

$$\begin{cases} \partial_t f + v \partial_x f + E \partial_v f = 0, & t > 0, \quad (x, v) \in I \times \mathbb{R}, \\ \lambda_D^2 \partial_x E = \int_{\mathbb{R}} f dv - n_e, & t > 0, \quad x \in I. \end{cases}$$

Here  $n_e$  is the density of electrons  $E = -\partial_x \log n_e$ , and  $\lambda_D > 0$  is the Debye length:  $n_e = \alpha^2 \exp \phi$  where  $\phi$  is the electric potential.

The interesting physical regime is when  $\lambda_D$  is small.

The formal limit  $\lambda_D \rightarrow 0^+$  equation has been called the Vlasov-Dirac equation

$$\partial_t f + v \partial_x f - \partial_x \log \left( \int_{\mathbb{R}} f dv \log \right) \partial_v f = 0, \quad t > 0, \quad (x, v) \in I \times \mathbb{R}$$

known to be singular or close to be singular.

- Bardos, Nouri: A Vlasov equation with Dirac potential used in fusion plasmas, private.
- Jabin, Nouri: Analytic solutions to a strongly nonlinear Vlasov equation, 2011.

Introduction

Moments

Stepping from  
Antonov to  
Morrison

Influence of the  
Debye length

Real non linear

Conclusion

A question about  
magnetic  
equation

Expansion  $f = G + g$ ,  $E = 0 + e$  and  $n_e = \alpha^2(1 + \phi + O(\phi^2))$  yields

$$\begin{cases} \partial_t g + v \partial_x g + e \partial_v G = 0, & t > 0, \quad (x, v) \in I \times \mathbb{R}, \\ -\lambda_D^2 \partial_{xx} \phi + \alpha^2 \phi = \int_{\mathbb{R}} g dv, & t > 0, \quad x \in I, \\ e = -\partial_x \phi, & t > 0, \quad x \in I. \end{cases}$$

The integro-differential operator becomes

$$\begin{aligned} L^D u(x, v) &= (-\lambda_D^2 \partial_{xx} + \alpha^2 + q(v)) u(x, v) \\ &\quad - v M(v) P.V. \int_{\mathbb{R}} \frac{1}{w - v} u(x, w) M(w) dw. \end{aligned}$$

- Much more in Morrison's papers

# Linear damping of the field

Introduction

Moments

Stepping from  
Antonov to  
Morrison

**Influence of the  
Debye length**

Real non linear

Conclusion

A question about  
magnetic  
equation

Assume that  $u_0 \in L^2(I : H^n(\mathbb{R}))$  with  $\underline{n \geq 2}$ .

Then the electric field tends to zero uniformly with respect to the Debye length with a rate at least  $t^{-n}$ .

Introduction

Moments

Stepping from Antonov to Morrison

Influence of the Debye length

Real non linear

Conclusion

A question about magnetic equation

The usual non linear equation

$$\mathbf{V.A.} \quad \begin{cases} \partial_t f + v \partial_x f - E \partial_v f = 0, & t > 0, \quad (x, v) \in I \times \mathbb{R}, \\ \partial_t E = \int_{\mathbb{R}} v f dv, & t > 0, \quad x \in I. \end{cases}$$

rewrites

$$\begin{cases} \partial_t u + v \partial_x u + v M E & = E \partial_v u - \frac{1}{2} v E u, \\ \partial_t E - \int u v M dv & = 0. \end{cases}$$

The quadratic energy-**Antonov**' energy equation is

$$\frac{d}{dt} \left( \int_I \int_{\mathbb{R}} u^2 dx dv + \int_I E^2 dx \right) = - \int_I \left( E \int_{\mathbb{R}} v u^2 dv \right) dx.$$

The  $E$  is "gentle" (since bounded in  $L^\infty$ ) by the usual theory.  
The  $v$  in  $\int_{\mathbb{R}} v u^2 dv$  is the dangerous term.

- Rein: Collisionless Kinetic Equations from Astrophysics—The Vlasov-Poisson System, 2007

Introduction

Moments

Stepping from  
Antonov to  
MorrisonInfluence of the  
Debye length

Real non linear

Conclusion

A question about  
magnetic  
equation

There is a "compatibility" between the left part of the system and its right part:

$$v\partial_x u + \frac{1}{2}vEu = ve^{\frac{1}{2}\phi}\partial_x\left(e^{-\frac{1}{2}\phi}u\right)$$

where  $\phi$  is the electric potential  $\partial_x\phi = -E$ .

# One more change of variable

Introduction

Moments

Stepping from Antonov to Morrison

Influence of the Debye length

Real non linear

Conclusion

A question about magnetic equation

Set

$$w(t, x, v) = e^{-\frac{1}{2}\phi(t,x)} u(t, x, v) \text{ and } F(t, x) = e^{-\frac{1}{2}\phi(t,x)} E(t, x)$$

which satisfy

$$\begin{cases} \partial_t w + v \partial_x w & + v M F & = E \partial_v w & - \frac{1}{2} \partial_t \phi w, \\ \partial_t F & - \int w v M dv & = & - \frac{1}{2} \partial_t \phi F. \end{cases}$$

The balance of the modified **Antonov**' energy writes

$$\frac{d}{dt} \mathcal{E}(t) = - \int \partial_t \phi \int w^2 dv dx - \int \partial_t \phi F^2 dx$$

where  $\mathcal{E} = \frac{1}{2} \int \int |w|^2 dx dv + \frac{1}{2} \int |F|^2 dx$ .

Introduction

Moments

Stepping from Antonov to Morrison

Influence of the Debye length

Real non linear

Conclusion

A question about magnetic equation

Only "gentle" terms in the RHS:  $\|\partial_t \phi(t)\|_{L^\infty(I)} \leq C \min\left(1, \mathcal{E}(t)^{\frac{1}{2}}\right)$   
 since

$$\partial_x(\partial_t \phi) = -\partial_t E = - \underbrace{\int_{\mathbb{R}} v f dv}_{\|\cdot\|_L^{\frac{3}{2}}(I) \leq C} = - \underbrace{\int_{\mathbb{R}} uv M dv}_{\|\cdot\|_L^2(I) \leq \mathcal{E}^{\frac{1}{2}}}.$$

So:  $\mathcal{E}' \leq C\mathcal{E} \times \min\left(1, \mathcal{E}^{\frac{1}{2}}\right)$ .

*Consider an initial data  $(u_0, E_0)$  (resp.  $(w_0, F_0)$ ) with finite quadratic norm. Then  $u(t)$  has finite quadratic norm for all time.*

*Moreover  $\mathcal{E}(t) \leq 2\mathcal{E}(0)$  for all  $t \leq T_{**} = O\left(\mathcal{E}(0)^{-\frac{1}{2}}\right)$ .*

In order words: if the **Antonov'** energy is small initially, it remains approximatively constant during a quite long time.

Introduction

Moments

Stepping from  
Antonov to  
Morrison

Influence of the  
Debye length

Real non linear

Conclusion

A question about  
magnetic  
equation

These tools can be used in some directions.

- Design test problems: low regularity initial data, preservation of quadratic norms, ...
- Truncate the system of moments and discretize it.
- The idea of moment has been discussed in some talks in the Workshop, for example for gyro-kinetic models and for the numerics.

- Helluy' talk: Numkin 2013.



In view of numerical modeling, start from a multiD **V.P.** model

$$\left\{ \begin{array}{l} \partial_t f_i + v \cdot \nabla_x f_i + \frac{q_i}{m_i} \nabla_v \cdot [(\mathbf{E} + v \wedge \mathbf{B}_{\text{ref}}) f_i] = 0, \\ \text{div} \mathbf{E} = \frac{1}{\epsilon_0} \left( q_i \int f_i(v) dv - n_e \right), \\ \mathbf{E} = -\frac{k_B T_e}{q} \nabla_x \ln(n_e) \end{array} \right.$$

or with its gyrokinetic approximation.

Assume  $\mathbf{B}_{\text{ref}}$  is to be modeled to study the coupling. Since

$$\text{curl} \mathbf{B} = \underbrace{\mu_0 (n_I \mathbf{u}_I - n_e \mathbf{u}_e)}_{\text{electric current}} + \underbrace{\mu_0 \epsilon_0}_{=1/c^2 \ll 1} \partial_t \mathbf{E}$$

$$\Rightarrow \mathbf{u}_e \approx \frac{1}{n_e} \left[ n_I \mathbf{u}_I - \frac{\text{curl} \mathbf{B}}{\mu_0 q} \right].$$

So one needs a companion equation which could be Ampère+Generalized Ohm's law.

Alternative in - Bonitto: Numkin 2013.

Introduction

Moments

Stepping from  
Antonov to  
Morrison

Influence of the  
Debye length

Real non linear

Conclusion

A question about  
magnetic  
equation

$$\partial_t \mathbf{B} + \text{curl} \mathbf{E} = 0$$

$$n_e \mathbf{E} = -\frac{k_B T_e}{e} \nabla(n_e) + \underbrace{\left[ \frac{\text{curl} \mathbf{B}}{\mu_0 q} - n_I \mathbf{u}_I \right]}_{\text{comes from } -u_e \wedge \mathbf{B}} \wedge \mathbf{B} + n_e \underbrace{\eta \text{curl} \mathbf{B}}_{\text{Joule effect}}$$

$$\left\{ \begin{array}{l} -\lambda^2 \Delta \ln n_e = n_I - n_e, \\ \frac{\partial \mathbf{B}}{\partial t} - \text{curl} \left( \frac{n_I \mathbf{u}_I}{n_e} \wedge \mathbf{B} \right) + \overbrace{\text{curl} \left( \frac{\text{curl} \mathbf{B}}{n_e} \wedge \mathbf{B} \right)}^{\text{Hall effect}} + \overbrace{\text{curl} (\eta \text{curl} \mathbf{B})}^{\text{Joule effect}} = 0, \\ \frac{\partial f}{\partial t} + \mathbf{v} \cdot \nabla f + \nabla_{\mathbf{v}} \cdot \left[ \left( -\nabla \ln n_e + \frac{\text{curl} \mathbf{B} - n_I \mathbf{u}_I}{n_e} \wedge \mathbf{B} + \mathbf{v} \wedge \mathbf{B} \right) f \right] = 0, \\ \text{div} \mathbf{B} = 0. \end{array} \right.$$

**Main point:** Correct energy balance and existence of weak solution of the problem with physically sound boundary conditions.

Question: it is reasonable from the physical viewpoint? At least  $\lambda_D$  small is necessary.

Can it be interesting for numerical models?

- Charles, D. Perthame, Sentis: Nonlinear stability of a Vlasov equation for magnetic plasmas, KRM 2013.

- Introduction
- Moments
- Stepping from Antonov to Morrison
- Influence of the Debye length
- Real non linear
- Conclusion
- A question about magnetic equation**

Thanks for your attention!